# Pressure drop due to the motion of a sphere near the wall bounding a Poiseuille flow 

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An expression is derived for the (low Reynolds number) additional pressure drop created by a relatively small sphere moving near the wall of a circular tube through which there is a Poiseuille flow. Two specific applications are examined: (i) the sedimentation of a homogeneous non-neutrally buoyant sphere in a quiescent fluid; and (ii) the motion of a neutrally buoyant sphere. In the latter case a pronounced increase in the additional pressure drop is predicted when the separation between the sphere and the tube wall is reduced to zero.

This analysis, which includes the behaviour for a sphere in contact with the tube wall, supplements previous 'method of reflexions' treatments valid only when the distance from the sphere centre to the wall is large compared with the sphere radius.

## 1. Introduction

One of the fundamental theoretical problems in the slow viscous duct flow of dilute, rigid-sphere suspensions is the description of the behaviour of a single sphere immersed in a fluid bounded by a circular cylindrical tube. The case of greatest interest in applications is that in which the sphere radius is small compared with that of the cylinder. This problem can be viewed as one of determining the hydrodynamic influence of the tube wall upon the suspended sphere. The magnitude of this interaction effect depends inter alia upon the size and lateral position of the sphere relative to the tube wall.

Prior analyses (based upon 'method of reflexions' or equivalent computational schemes) have been concerned exclusively with the case where the distance of the sphere from the tube wall is large compared with the sphere radius. The present work is devoted to a treatment of the opposite case, where the sphere is situated very near to the wall. It will be demonstrated that very considerable alterations can occur in the translational and rotational particle velocities, and in the additional pressure drop (above the purely Poiseuillian portion of the pressure drop) when a small sphere is in close proximity to the wall, in contrast to the case where it is relatively distinct. The consequences of this hydrodynamic wall effect are pertinent to a proper understanding of suspension sedimentation and rheology.

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Consider a homogeneous sphere (radius $=a$, density $=\rho_{s}$ ) immersed in an incompressible Newtonian fluid (viscosity $=\mu$, density $=\rho$ ) confined within a long circular cylindrical tube of radius $R_{o}$. When the sphere is not neutrally buoyant ( $\rho_{s} \neq \rho$ ) the tube will be assumed vertical, with its symmetry axis parallel to gravity. As in figure 1 , let $h$ denote the radial distance from the sphere centre to the nearest point on the tube wall.

All prior theoretical treatments of flow past an eccentrically positioned sphere have been concerned with the limit in which the sphere is small compared with its distance from the wall, i.e. $a \ll h \leqslant R_{o}$. As shown by Cox \& Brenner (1967) the slow viscous motion in this limit can be analysed via regular perturbation expansion procedures. Alternatively, equivalent asymptotic expansions may be derived by application of the 'method of reflexions'. This latter technique was applied by Brenner \& Happel (1958) to Poiseuille flow past a sphere translating, without rotation, parallel to the tube axis. Higher order terms in $a / R_{o}$ were subsequently evaluated by Greenstein \& Happel (1968, 1970), who also included the effect of sphere rotation. The results of Greenstein \& Happel (1970) provide an asymptotic expression for the additional pressure drop in the important special case in which sedimentation and other buoyancy effects are negligible. Concurrently, Brenner (1970) demonstrated that the same neutrally buoyant pressure drop terms could be obtained without the detailed knowledge of the boundary-value solutions required by the method of reflexions. When the sphere is constrained to translate (without rotation) along the tube axis ( $h=R_{o}$ ) the fluid motion is axisymmetric, and hence easily susceptible to higher order analysis (Haberman \& Sayre 1958; Hochmuth \& Sutera 1970; Wang \& Skalak 1969).

Because of convergence difficulties the method of reflexions, as well as equivalent expansion procedures, cannot be used to examine the hydrodynamic particlewall interaction experienced by an eccentrically positioned sphere in the case of either ( $a$ ) a small sphere in close proximity to the tube wall, $a \leqslant h \ll R_{o}$; or (b) a closely fitting sphere, $a \leqslant h \leqslant R_{o}, a=O\left(R_{o}\right)$. The proper analysis of case (b), which necessitates employing singular perturbation techniques, is treated in a companion paper (Bungay \& Brenner 1973b; see also Bungay 1970). In the present study an alternative regular perturbation procedure is outlined for case (a) to complement the method of reflexions.

Before confining attention to this case we shall derive reciprocity relationships between the additional pressure drop and the hydrodynamic force and torque acting on a sphere. These apply without restriction as to sphere/tube size or radial position in the tube.

## 2. Sphere of arbitrary size and position

## Formulation of the problem

In figure 1 a Cartesian co-ordinate system $(x, y, z)$ is defined with origin at the centre $o$ of the sphere. The unit vector $\hat{\mathbf{z}}$ of the co-ordinate system is parallel to the tube axis. Both the $\hat{\mathbf{y}}$ and $\hat{\mathbf{z}}$ unit vectors lie in the meridian plane containing


Figure 1. Eccentrically positioned sphere in a vertical cylindrical tube.
the tube axis and passing through the sphere centre. The cross-sectional plane perpendicular to the tube axis and passing through the point $o$ constitutes a second plane of geometrical reflexion symmetry.

The suspending fluid flows upward through the vertical tube at a steady superficial mean velocity $V_{m}$. Motion of the sphere is prescribed by the translational velocity $\mathbf{U}_{o}$ of its centre and its angular velocity $\boldsymbol{\omega}$. Accordingly, in consequence of the no-slip condition, the local fluid motion satisfies the boundary condition

$$
\begin{equation*}
\mathbf{v}=\mathbf{U}_{o}+\boldsymbol{\omega} \times \mathbf{r} \quad \text { on the sphere }, \tag{2.1}
\end{equation*}
$$

$$
\begin{equation*}
\mathbf{v}=\mathbf{0} \quad \text { on the tube wall, } \tag{2.2}
\end{equation*}
$$

where $\mathbf{r}$ is the position vector relative to an origin at the sphere centre. The three characteristic velocities $\mathbf{U}_{o}, \boldsymbol{\omega}$ and $V_{m}$ may be tentatively regarded as independent quantities. In consequence of these motions the particle is acted upon by a hydrodynamic force $\mathbf{F}$ and a hydrodynamic torque $\mathbf{T}_{o}$ about the sphere centre.

All relevant particle Reynolds numbers are assumed sufficiently small compared with unity to justify neglect of the inertial terms in the Navier-Stokes equations. After the hydrostatic term has been incorporated in the definition of the dynamic pressure field $p$, the dynamical and kinematical equations governing the fluid motion thereby reduce to the Stokes equations,
and

$$
\begin{align*}
\mu \nabla^{2} \mathbf{v} & =\nabla p  \tag{2.3}\\
\nabla \cdot \mathbf{v} & =0 . \tag{2.4}
\end{align*}
$$

As demonstrated by Cox \& Mason (1971) for the case where gravity represents the only external force, the sphere cannot experience any lateral movement. Rather, the sphere, if translating, necessarily travels parallel to the tube axis. Hence, we may write

$$
\begin{equation*}
\mathbf{U}_{o}=U_{o} \hat{\mathbf{z}}, \quad \mathbf{F}=F \hat{z} \tag{2.5a,b}
\end{equation*}
$$

By symmetry arguments based upon the sphere-tube geometry, the translational motion (2.5a) can give rise to a hydrodynamic torque possessing a component only in the $x$ direction. If external torques other than those acting parallel to $\hat{\mathbf{x}}$ are supposed absent in general, then we may also prescribe that

$$
\begin{equation*}
\omega=\omega \hat{\mathbf{x}}, \quad \mathbf{T}_{o}=T_{o} \hat{\mathbf{x}} . \tag{2.6a,b}
\end{equation*}
$$

The disturbance to the Poiseuille flow caused by the presence of the sphere decays exponentially with distance from the particle (Sonshine, Cox \& Brenner 1966). Consider a hypothetical 'inlet' plane $S_{i}$ located at a distance of $\frac{1}{2} l$ units upstream of the sphere and an 'exit' plane $S_{e}$ at a similar distance downstream. For sufficiently large $l$ the local velocity field is Poiseuillian:

$$
\begin{gather*}
\mathbf{v}\left(S_{i}\right), \mathbf{v}\left(S_{e}\right) \rightarrow u \hat{\mathbf{z}} \quad \text { as } \quad l \rightarrow \infty,  \tag{2.7}\\
u=2 V_{m}\left[1-\left(R / R_{o}\right)^{2}\right], \tag{2.8}
\end{gather*}
$$

with
in which $R$ is the radial distance from the tube axis. The dynamic pressure across each of these two distant planes is uniform, and consequently can be written as

$$
\left.\begin{array}{rl}
p\left(S_{i}\right) & \stackrel{*}{p}\left(S_{i}\right)+\frac{1}{2} \Delta P^{+}  \tag{2.9}\\
p\left(S_{e}\right) \rightarrow \stackrel{*}{p}\left(S_{e}\right)-\frac{1}{2} \Delta P^{+}
\end{array}\right\} \quad \text { as } \quad l \rightarrow \infty
$$

where $\stackrel{*}{p}=-8 \mu V_{m} z / R_{o}^{2}+C$ is the Poiseuille pressure distribution for flow at mean velocity $V_{m}$ in the absence of the sphere, $C$ being an arbitrary constant. Let
and

$$
\begin{equation*}
\Delta P=p\left(S_{i}\right)-p\left(S_{e}\right) \tag{2.10}
\end{equation*}
$$

$$
\begin{equation*}
\Delta \stackrel{*}{P}=\stackrel{*}{p}\left(S_{i}\right)-\stackrel{*}{p}\left(S_{e}\right), \tag{2.11}
\end{equation*}
$$

respectively, denote the pressure drops in the presence and absence of the particle from the Poiseuille flow. Then,

$$
\begin{equation*}
\Delta P^{+}=\lim _{l \rightarrow \infty}(\Delta P-\Delta \stackrel{*}{P}) \tag{2.12}
\end{equation*}
$$

represents the additional pressure drop due to the presence of the sphere in the flow.

In view of the linearity of the Stokes equations and of the boundary conditions, the hydrodynamic force, torque and additional pressure drop are necessarily linear functions of each of the three characteristic velocities. Thus, we may write

$$
\left(\begin{array}{c}
F  \tag{2.13}\\
T_{o} \\
\Delta P+A
\end{array}\right)=-\mu\left(\begin{array}{lll}
K^{t} & K^{r} & K^{s} \\
L^{t} & L^{r} & L^{s} \\
P^{t} & P^{r} & P^{s}
\end{array}\right)\left(\begin{array}{c}
U_{o} \\
\omega \\
-V_{m}
\end{array}\right)
$$

in which $A=\pi R_{o}^{2}$ is the tube cross-sectional area, $\Delta P^{+} A$ then being the additional pressure drop force. The matrix composed of the nine 'intrinsic hydrodynamic resistance' coefficients is a purely geometrical function, dependent only upon the sphere and tube radii, and upon the relative radial position of the sphere in the tube.

The scalar elements of the hydrodynamic resistance matrix are all positive, with the exception of $K^{r}$ and $L^{t}$. In the perturbed motion, representing the difference between the flows in the presence and absence of the particle, the rate $\dot{W}$ at which work is being done by the stresses acting over the surfaces bounding the fluid is $\dot{W}=-\left(F U_{o}+T_{o} \omega-\Delta P^{+} A V_{m}\right)$. In creeping flow this quantity is identical to the rate of mechanical energy dissipation $E$ in the fluid bounded by the surfaces $S_{i}$ and $S_{e}$, the tube wall $S_{w}$ and the sphere surface $S_{p}$. Since this dissipation rate is necessarily non-negative it follows that

$$
\dot{E}=-\left(U_{o}, \omega,-V_{m}\right)\left(\begin{array}{c}
F  \tag{2.14}\\
T_{o} \\
\Delta P^{+} A
\end{array}\right) \geqslant 0
$$

whence the resistance matrix is a positive-definite form.
The resistance matrix also enjoys the property of being symmetric, corresponding to the three reciprocity relations

$$
\begin{equation*}
K^{r}=L^{t}, \quad P^{t}=K^{s}, \quad P^{r}=L^{s} \tag{2.15}
\end{equation*}
$$

The first of these is proved by Brenner (1964). Demonstrations of the validity of the remaining two Onsager-like reciprocal relations are furnished in the following paragraphs. Equations (2.13)-(2.17) apply even for non-circular cylindrical ducts.

## Reciprocal relations

Derivations of (2.16) and (2.17) may be formulated via the Lorentz reciprocal theorem using techniques previously applied in the proof of (2.15). Let $\mathbf{v}^{\prime}$ and $v^{\prime \prime}$ be any pair of velocity fields satisfying the Stokes equations, and let $\pi^{\prime}$ and $\pi^{\prime \prime}$ denote the corresponding pressure tensors: $\boldsymbol{\pi}=-\mathbf{I} p+\mu\left[\nabla \mathbf{v}+(\nabla \mathbf{v})^{\dagger}\right]$. According to the reciprocal theorem,

$$
\begin{equation*}
\oint_{S} d \mathbf{S} \cdot \boldsymbol{\pi}^{\prime} \cdot \mathbf{v}^{\prime \prime}=\oint_{S} d \mathbf{S} \cdot \boldsymbol{\pi}^{\prime \prime} \cdot \mathbf{v}^{\prime} . \tag{2.18}
\end{equation*}
$$

Here, $S$ is an arbitrary closed surface bounding a continuous volume $V$ of the homogeneous fluid, and $d \mathbf{S}$ is a differential element of surface area directed into the fluid volume $V$. For present purposes, $S$ is chosen to consist of the planes $S_{e}$ and $S_{i}$, the portion $S_{w}$ of the inside wall of the tube between these two planes, and the surface $S_{p}$ of the sphere

$$
\begin{equation*}
S=S_{e}+S_{i}+S_{p}+S_{w} \tag{2.19}
\end{equation*}
$$

To deduce the cross-relationship (2.16), let the primed motion denote flow past a sphere which is stationary with respect to the tube. Thus, at the rigid surfaces,

$$
\begin{equation*}
\mathbf{v}^{\prime}=\mathbf{0} \quad \text { on } \quad S \text { and } S_{w} \tag{2.20a,b}
\end{equation*}
$$

while up- and downstream of the sphere,

$$
\begin{equation*}
\mathbf{v}^{\prime} \rightarrow u \hat{\mathbf{z}} \quad \text { on } \quad S_{e} \text { and } S_{i} . \tag{2.20c}
\end{equation*}
$$

Choice of the double-primed flow to be that resulting from the sphere translating, without rotation, through an otherwise quiescent medium ( $V_{m}=0$ ) leads to the boundary conditions

$$
\begin{align*}
\mathbf{v}^{\prime \prime}= & U_{0} \hat{\mathbf{z}} \quad \text { on } \quad S_{p}, \quad \mathbf{v}^{\prime \prime}=\mathbf{0} \quad \text { on } \quad S_{w}  \tag{2.21a,b}\\
& \mathbf{v}^{\prime \prime} \rightarrow \mathbf{0} \quad \text { on } \quad S_{e} \text { and } S_{i} . \tag{2.21c}
\end{align*}
$$

and
From the last of these conditions it follows that $\pi^{\prime \prime} \rightarrow \mathbf{I} p^{\prime \prime}$, with $p^{\prime \prime}$ being uniform on both $S_{e}$ and $S_{i}$. This limit condition, together with ( $2.20 a, b, c$ ), when substituted into the right-hand side of (2.18), along with $d \mathbf{S}=+\hat{\mathbf{z}} d S$ on $S_{i}$ and

$$
d \mathbf{S}=-\hat{\mathbf{z}} d S \quad \text { on } \quad S_{e},
$$

yields

$$
\begin{equation*}
\oint_{S} d \mathbf{S} \cdot \pi^{\prime \prime} \cdot \mathbf{v}^{\prime}=\left[p^{\prime \prime}\left(S_{e}\right)-p^{\prime \prime}\left(S_{i}\right)\right] \int_{S_{i}} u d S \rightarrow \mu P^{t} U_{o} V_{m} \tag{2.22}
\end{equation*}
$$

In arriving at the above, (2.12) and (2.13) have been used along with

$$
V_{m}=A^{-1} \int_{S_{i}} u d S \equiv A^{-1} \int_{S_{e}} u d S
$$

Application of (2.5b) and (2.13), and use of the integral relation

$$
\mathbf{F}=\int_{S_{p}} d \mathbf{S} . \boldsymbol{\pi}
$$

to evaluate the hydrodynamic force exerted on the sphere in the single-primed case, leads to the relation

$$
\int_{S_{p}} d \mathbf{S} \cdot \boldsymbol{\pi}^{\prime} \cdot \hat{\mathbf{z}}=\mu K^{s} V_{m}
$$

This expression, together with boundary conditions (2.21a,b,c), simplifies the left-hand side of (2.18) to the form

$$
\begin{equation*}
\oint_{S} d \mathbf{S} \cdot \pi^{\prime} \cdot \mathbf{v}^{\prime \prime}=\mu K^{s} U_{o} V_{m} \tag{2.23}
\end{equation*}
$$

Identity (2.16) is established by equating (2.22) and (2.23). The second identity, (2.17), is proved in the same manner upon choosing the double-primed flow to be that arising solely from rotation of the sphere ( $U_{0}=V_{m}=0$ ).

The reciprocity relation analogous to (2.16) for the motion of a spherical fluid droplet within a duct of arbitrary cross-section was derived by Bungay \& Brenner (1973a).

## 3. Small sphere near the tube wall

Perturbation expansion
Consider the case of a sphere whose relative size and lateral position in the tube satisfy the dual constraints
in which

$$
\begin{gather*}
\lambda \ll a / h \leqslant 1,  \tag{3.1}\\
\lambda=a / R_{o} . \tag{3.2}
\end{gather*}
$$

Expressed in the Cartesian co-ordinates of figure 1, the equation of the tube wall surface is $x^{2}+\left(R_{o}-h-y\right)^{2}=R_{o}^{2}$. Use of the dimensionless co-ordinates

$$
\begin{equation*}
\hat{x}=x / a, \quad \hat{y}=y / a, \quad \hat{z}=z / a \tag{3.3}
\end{equation*}
$$

permits the equation describing the wall to be expressed as a series expansion in powers of $\lambda$ :

$$
\begin{equation*}
\hat{y}=-h / a+\left(\frac{1}{2} \hat{x}^{2}\right) \lambda+O\left(\lambda^{2}\right) \tag{3.4}
\end{equation*}
$$

for $|\hat{x}| \ll \lambda^{-1}$. This expansion shows that, to terms of zero order in $\lambda$, the tube wall may be replaced by the plane surface $\hat{y}=-h / a$.

The asymptotic expression (3.4) for the surface defining the cylindrical boundary suggests that the velocity and pressure fields possess the following regular perturbation expansions:
and

$$
\begin{align*}
\mathbf{v} & =V_{*}\left[\mathbf{v}_{0}+\lambda \mathbf{v}_{\mathbf{1}}+\ldots\right]  \tag{3.5}\\
p & =\left(\mu V_{*} / a\right)\left[p_{0}+\lambda p_{1}+\ldots\right] \tag{3.6}
\end{align*}
$$

in which $V_{*}$ is a characteristic reference velocity to be determined. The local field pairs $\left(\mathrm{v}_{0}, p_{0}\right),\left(\mathbf{v}_{1}, p_{1}\right), \ldots$, depend upon $\hat{\mathbf{r}}=\mathbf{r} / a$ and the parameter $a / h$, but are independent of $\lambda$ for $\lambda \ll a / h$. Upon substitution of the expansions (3.5) and (3.6) into the governing equations of $\S 2$, the zero-order fields ( $\mathrm{v}_{0}, p_{0}$ ) are found to satisfy the following differential equations and boundary conditions:

$$
\begin{gather*}
\hat{\nabla}^{2} \mathbf{v}_{0}=\hat{\nabla} p_{0}, \quad \hat{\nabla} \cdot \mathbf{v}_{0}=0  \tag{3.7a,b}\\
\mathbf{v}_{0}=\hat{U_{o}} \hat{\mathbf{z}}+\hat{\omega}(\hat{\mathbf{x}} \times \hat{\mathbf{r}}) \quad \text { on the sphere }|\hat{\mathbf{r}}|=1,  \tag{3.8}\\
\mathbf{v}_{0}=\mathbf{0} \quad \text { on the plane } \hat{y}=-h / a  \tag{3.9}\\
\mathbf{v}_{\mathbf{0}} \rightarrow \mathbf{u}_{0}, \quad p_{0} \rightarrow \mathbf{0} \quad \text { as }|\hat{\mathbf{r}}| \rightarrow \infty  \tag{3.10}\\
\hat{U}_{o}=U_{o} / V_{*}, \quad \hat{\omega}=\omega a / V_{*}  \tag{3.11a,b}\\
\hat{\nabla}=a \nabla \tag{3.12}
\end{gather*}
$$

and
in which
and
The zero-order undisturbed velocity field $\mathbf{u}_{0}$ far from the sphere can be found by expressing the Poiseuille velocity distribution as a polynomial in $\lambda$ of the form

$$
\begin{equation*}
\mathbf{u}=V_{*}\left[\mathbf{u}_{0}+\lambda \mathbf{u}_{\mathbf{1}}+\ldots\right] \tag{3.13}
\end{equation*}
$$

A concise formulation of the zero-order field found in this way is
in which

$$
\begin{equation*}
\mathbf{u}_{0}=\hat{\mathbf{u}}_{0}+\hat{\mathbf{r}} \cdot(\hat{\nabla} \hat{\mathbf{u}})_{o} \tag{3.14}
\end{equation*}
$$

and

$$
\begin{equation*}
(\hat{\nabla} \hat{\mathbf{u}})_{o}=(a / h) \hat{\mathbf{y}} \hat{\mathbf{z}} . \tag{3.15}
\end{equation*}
$$

This recasting of the Poiseuille field also shows that the reference velocity should be chosen as

$$
\begin{equation*}
V_{*}=4 V_{m} h / R_{o} . \tag{3.17}
\end{equation*}
$$

Equation (3.14) constitutes a simple shear flow whose dimensional rate of shear $4 V_{m} / R_{o}$ is the local Poiseuille shear rate at the tube wall. Equations (3.7)-(3.10) and (3.14) reveal that, to terms of zero order in $\lambda$, the original tube flow problem may be approximated by the more tractable problem of simple shear flow round a sphere in the neighbourhood of a plane wall.

Expansions (3.5) and (3.6) induce expansions of the hydrodynamic force and torque of the forms
and

$$
\begin{align*}
& F=6 \pi \mu V_{*} a\left[F_{0}+\lambda F_{1}+\ldots\right]  \tag{3.18}\\
& T_{o}=8 \pi \mu V_{*} a^{2}\left[\left(T_{o}\right)_{0}+\lambda\left(T_{o}^{\prime}\right)_{1}+\ldots\right] \tag{3.19}
\end{align*}
$$

The zero-order contributions expressed in terms of dimensionless zero-order force and torque resistance coefficients are

$$
\begin{align*}
F_{0} & =K_{0}^{s}-K_{0}^{t} \hat{U}_{o}-K_{0}^{r} \hat{\omega}  \tag{3.20}\\
\left(T_{o}\right)_{0} & =L_{0}^{s}-L_{0}^{t} \widehat{U}_{o}-L_{0}^{r} \hat{\omega} . \tag{3.21}
\end{align*}
$$

and
There exist no zero-order contributions to the pressure drop coefficients in (2.13). This is to be expected since no pressure drop can arise from the semiinfinite flow described by (3.7)-(3.10). However, the zero-order flow fields do give rise to pressure drops of second order in $\lambda$. Through application of the reciprocal theorem in a derivation similar to those of $\S 2$, Brenner (1970) has shown that the additional pressure drop for a rigid particle in creeping duct flow can be calculated from the expression

$$
\begin{equation*}
\Delta P^{+} V_{m} A=\int_{S_{p}} d \mathbf{S} . \boldsymbol{\pi} \cdot \mathbf{u} \tag{3.22}
\end{equation*}
$$

The pressure tensor expanded in powers of $\lambda$ is

$$
\begin{equation*}
\pi=\left(\mu V_{*} / a\right)\left[\pi_{0}+\lambda \pi_{1}+\ldots\right] \tag{3.23}
\end{equation*}
$$

Substituting expansions (3.13) and (3.23) into (3.22), along with use of $d \mathbf{S}=a^{2} d \hat{\mathbf{S}}$, yields

$$
\begin{equation*}
\Delta P^{+}=\left(\mu V_{*} / R_{o}\right)\left[\lambda^{2} \Delta P_{2}^{+}+\lambda^{3} \Delta P_{3}^{+}+\ldots\right], \tag{3.24}
\end{equation*}
$$

where, for the leading coefficient,

$$
\begin{equation*}
\Delta P_{2}^{+}=\frac{4}{\pi}\left(\frac{h}{a}\right) \int_{S_{p}} d \hat{\mathbf{S}} \cdot \boldsymbol{\pi}_{0} \cdot \mathbf{u}_{0} \tag{3.25}
\end{equation*}
$$

in which $\pi_{0}=-I p_{0}+\hat{\nabla} \mathbf{v}_{\mathbf{0}}+\left(\hat{\nabla} \mathbf{v}_{\mathbf{0}}\right)^{\dagger}$ is the dimensionless pressure tensor derived from the solution ( $\mathbf{v}_{0}, p_{0}$ ) of (3.7)-(3.10). This coefficient may be expressed linearly in terms of dimensionless second-order resistance coefficients as

$$
\begin{equation*}
\Delta P_{2}^{+}=P_{2}^{\mathrm{s}}-P_{2}^{\mathrm{t}} \hat{U}_{o}-P_{2}^{r} \hat{\omega} \tag{3.26}
\end{equation*}
$$

The dimensionless reciprocity relations corresponding to (2.15)-(2.17) for the linear shear flow problem are then

$$
\begin{gather*}
K_{0}^{r}=\frac{4}{3} L_{0}^{t}  \tag{3.27}\\
P_{2}^{t}=24(h / a) K_{0}^{s}, \quad P_{2}^{r}=32(h / a) L_{0}^{s} \tag{3.28}
\end{gather*}
$$

These resistance coefficients are dependent only on $a / h$, being independent of $\lambda$. Their variation with the parameter $a / h$ is discussed in the following subsection.

## Evaluation of resistance coefficients

The motion of a sphere in the neighbourhood of a plane wall has been studied extensively. Goldman, Cox \& Brenner (1967a,b) tabulate numerical values of the force and torque resistance coefficients over the full range of the parameter,

| $\alpha$ | $a / h$ | $P_{2}$ | $P_{2}^{s}$ | $P_{2}^{t}$ |
| :---: | :---: | :---: | :---: | :---: |
| $\infty$ | 0 | $16 \dagger$ | $\infty \dagger$ | $\infty \dagger$ |
| $3 \cdot 0$ | 0.09933 | 15.9970 | 257.859 | 255.817 |
| $2 \cdot 0$ | 0.26580 | 15.9538 | $110 \cdot 498$ | $105 \cdot 382$ |
| 1.5 | 0.42510 | 15.8416 | $79 \cdot 864$ | $72 \cdot 151$ |
| 1.0 | 0.64805 | $15 \cdot 5870$ | $64 \cdot 273$ | 53.297 |
| $0 \cdot 5$ | 0.88684 | 15.2598 | 57.921 | $43 \cdot 731$ |
| $0 \cdot 3$ | 0.95666 | $15 \cdot 1630$ | 56.944 | 41.849 |
| $0 \cdot 1$ | 0.99502 | 15-1107 | 56.516 | 40.929 |
| 0.05 | 0.99875 | 15.1056 | 56.477 | $40 \cdot 844$ |
| 0.01 | 0.99995 | $15 \cdot 1040$ | $56 \cdot 466$ | 40.814 |
| 0 | $1 \cdot 00000$ | 15.1038 | 56.465 | $40 \cdot 812$ |

Table 1. Pressure drop resistance coefficients. The bipolar co-ordinate parameter $\alpha$ is defined as $a / h=\operatorname{sech} \alpha$
$0<a / h \leqslant 1$. These values were computed from exact solutions of the Stokes equations in spherical bipolar co-ordinates, and from asymptotic singular perturbation analyses. The notation of these authors is related to that of the present text by the relations

$$
\begin{array}{ll}
K_{0}^{r}=-F_{x}^{\tau^{*}}, & L_{0}^{r}=-T_{y}^{r^{*}} \\
K_{0}^{s}=F_{x}^{s^{*}}, & L_{0}^{s}=\frac{1}{2}(a / h) T_{y}^{s^{*}} \\
K_{0}^{t}=-F_{x}^{t^{*}}, & L_{0}^{t}=-T_{y}^{t *} \tag{3.32a,b}
\end{array}
$$

Values of $K_{0}^{t}$ and $L_{0}^{t}$ had previously been calculated by O'Neill (1964), who also computed $K_{0}^{s}$ and $L_{0}^{s}$ for $a / h=1$ ( $O^{\prime}$ Neill 1968).

Values of the pressure drop coefficients $P_{2}^{r}$ and $P_{2}^{t}$ presented in table 1 were calculated from the results of Goldman et al. (1967b) by employing (3.31) in conjunction with the identities (3.28) and (3.29). $\dagger$ Computation of values of the remaining coefficient $P_{2}^{s}$ appearing in table 1 required use of the following integral expression, obtained from (3.25) and (3.26):

$$
\begin{equation*}
P_{2}^{\mathrm{s}}=\frac{4}{\pi}\left(\frac{h}{a}\right) \int_{S_{p}} d \hat{\mathbf{S}} \cdot \pi_{0}^{\mathrm{s}} \cdot \mathbf{u}_{0} \tag{3.33}
\end{equation*}
$$

The dimensionless pressure tensor $\pi_{0}^{8}$ appearing in the above integrand is that arising from the solution $\left(v_{0}, p_{0}\right)$ of (3.7)-(3.10) for a stationary sphere,

$$
\hat{U}_{0}=\hat{\omega}=0
$$

$\dagger$ The asymptotic forms of the pressure drop coefficients as $a / h \rightarrow 0$ are as follows:

$$
\begin{aligned}
& P_{2}^{r}=16\left[1-\frac{3}{16}\left(\frac{a}{h}\right)^{3}\right]+O\left(\frac{a}{h}\right)^{4}, \\
& P_{2}^{t}=24\left(\frac{h}{a}\right)\left[1-\frac{5}{16}\left(\frac{a}{h}\right)^{3}\right]\left[1-\frac{9}{16}\left(\frac{a}{h}\right)+\frac{1}{8}\left(\frac{a}{h}\right)^{3}-\frac{45}{256}\left(\frac{a}{h}\right)^{4}\right]^{-1}+O\left(\frac{a}{h}\right)^{4}, \\
& P_{2}^{s}=24\left(\frac{h}{a}\right)\left[1+\frac{8}{9}\left(\frac{a}{h}\right)^{2}-\frac{5}{8}\left(\frac{a}{h}\right)^{8}\right]\left[1-\frac{9}{16}\left(\frac{a}{h}\right)+\frac{1}{8}\left(\frac{a}{h}\right)^{3}\right]^{-1}+O\left(\frac{a}{h}\right)^{2}
\end{aligned}
$$

These reflexion-type expansions were obtained from the comparable reflexion-type explessions for the force and torque resistance coefficients summarized by Goldman et al. (1967a, b), the analysis of Wakiya, Darabaner \& Mason (1967), and the limiting form of the additional. pressure drop for a neutrally buoyant sphere suspended in a Poiseuille flow (Brenner 1970), $\Delta P_{2}^{+}=\frac{40}{3}(a / h)+O(a / h)^{2}$, derived from (3.24), (3.17) and (4.14) in the limit where $\beta \rightarrow 1$.

The exact bipolar co-ordinate solution of this problem (Goren \& O'Neill 1971) was used to effect an analytic integration of (3.34), using techniques developed by Goldman (1966) for the evaluation of a similar integral (Goldman et al. 1967b). $\dagger$

## 4. Results and discussion

The asymptotic expressions of $\S 3$ will be applied in two specific situations of importance: the sedimentation of a homogeneous non-neutrally buoyant sphere in an otherwise quiescent fluid; and the motion of a neutrally buoyant sphere suspended in a Poiseuille flow.

## Sedimentation of a homogeneous sphere in a quiescent fluid ( $V_{m}=0$ )

A sphere settling at its terminal speed is acted upon by a hydrodynamic force which is equal in magnitude, but opposite in direction, to the force of gravity $\mathbf{g}$ (corrected for the buoyant effect of the fluid), $F=\frac{4}{3} \pi a^{3}\left(\rho_{s}-\rho\right) g$. As the sphere is homogeneous, no external gravitational couple acts on it (i.e., $\Gamma_{o}=0$ ), whence the sphere is freely rotating. According to (2.13) the angular velocity of such a sphere is related to its terminal translational velocity by the expression

$$
\begin{equation*}
\omega=-\left(L^{t} / L^{r}\right) U_{o} . \tag{4.1}
\end{equation*}
$$

In the absence of the constraining effect of the tube walls the sphere translates without rotation, its sedimentation velocity being given by Stokes' law as

$$
\begin{equation*}
U_{\infty}=-F / 6 \pi \mu a \tag{4.2}
\end{equation*}
$$

On substituting (4.1) and (4.2) into (2.13) one finds, in comparison with the unbounded case, that the sphere in the tube settles at a velocity $U_{o}$ given by the relation

$$
\begin{equation*}
U_{o} / U_{\infty}=6 \pi a\left[K^{t}-\left(L^{t} / L^{r}\right) K^{r}\right]^{-1} . \tag{4.3}
\end{equation*}
$$

Sedimentation creates a pressure differential $\Delta P^{+}$in the fluid (Feldman \& Brenner 1968), the pressure in the fluid being highest on that side of the sphere towards which it translates. From (2.13) and (4.1) this pressure diminution is

$$
\Delta P^{+} A / \mu U_{o}=\left(L^{t} / L^{r}\right) P^{r}-P^{t}
$$

Alternatively, use of (2.13) in conjunction with the reciprocity conditions (2.15)-(2.17) gives the equivalent formula

$$
\begin{equation*}
\frac{\Delta P^{+} A}{F}=\frac{K^{s} L^{r}-K^{r} L^{s}}{\bar{K}^{t} L^{r}-\bar{K}^{r} L^{t}} \tag{4.4}
\end{equation*}
$$

Relations (4.3) and (4.4) apply for a sphere of any size or position in the tube. For a small sphere ( $\lambda \ll 1$ ) they adopt the asymptotic (dimensionless) forms
and

$$
\begin{gather*}
U_{o} / U_{\infty}=\left[K_{0}^{t}-\left(L_{0}^{t} / L_{0}^{r}\right) K_{0}^{r}\right]^{-1}+O(\lambda)  \tag{4.5}\\
\frac{\Delta P^{+} A}{F}=4\left(\frac{h}{a}\right) \frac{K_{0}^{s} L_{0}^{r}-K_{0}^{r} L_{0}^{s}}{K_{0}^{t} L_{0}^{r}-K_{0}^{r} L_{0}^{t}} \lambda+O\left(\lambda^{2}\right) \tag{4.6}
\end{gather*}
$$

These forms are valid for expansions of the type discussed in §3 $(\lambda \ll a / h \leqslant 1)$, as well as for 'reflexion-type' expansions ( $\lambda \leqslant a / h \ll 1$ ).

[^1]

Frgure 2. Terminal sedimentation velocity of a sphere settling freely in a vertical tube $\left(T_{o}=0, V_{m}=0\right)$ : ———, perturbation expansion (4.5); ----, method of reflexions expansion (4.7); O, sphere centre on tube axis.

The effect of the wall on the sedimentation velocity of a small sphere is illustrated in figure 2. The dashed lines are obtained from the method of reflexions expansion (Brenner 1966)

$$
\begin{equation*}
U_{o} / U_{\infty}=1-f(\beta) \lambda+O\left(\lambda^{3}\right) \quad(\beta \leqslant 1-\lambda), \tag{4.7}
\end{equation*}
$$

in which the argument $\beta$ of the wall-effect function $f$ represents the fractional radial distance of the sphere centre from the tube axis:

$$
\begin{equation*}
\beta=1-(h / a) \lambda . \tag{4.8}
\end{equation*}
$$

Numerical tabulations of the function $f(\beta) v s . \beta$ in the range $0 \leqslant \beta \leqslant 1$ are furnished by Greenstein \& Happel (1968) and, more completely, by Hirschfeld (1972). The values of $f(\beta)$ decrease slightly from $f(0)=2 \cdot 10444$ to a minimum at $\beta \approx 0.41$ before approaching the asymptote $f \sim(9 / 16)(1-\beta)^{-1}$ as $\beta \rightarrow 1$. The circled points from which the curves originate denote the velocity ratios for sedimentation along the tube axis. The solid line in figure 2 is a plot of the first term of expansion (4.5), obtained using values of the resistance coefficients tabulated by Goldman et al. (1967a). The solid and dashed lines will tend to agree for the intermediate parameter range $\lambda \ll a / h \ll 1$ common to both expansion procedures.

Not surprisingly, the velocity ratio $U_{o} / U_{\infty}$ never exceeds unity, the influence of the cylindrical boundary being to retard the settling motion of the sphere at all radial positions. Provided the constraint $\lambda \ll 1-\beta$ is satisfied the settling velocity for a given $\lambda$ is greatest at the lateral position ( $\beta \approx 0.41$ ) corresponding to the minimum in $f(\beta)$. For $\beta>0 \cdot 41$, the closer the sphere is to the wall the more slowly it settles. In fact, according to the solid curve in figure 2, the settling.


Figure 3. Pressure drop created by sphere settling in a vertical tube ( $T_{0}=0, V_{m}=0$ ) or translational velocity of a neutrally buoyant sphere ( $F=0, T_{o}=0$ ): ——, perturbation expansion (4.6); ----, method of reflexions expansion (4.9); $O$, sphere centre on tube axis; - data of Goldsmith \& Mason (1962).
velocity decreases rapidly to zero as $a / h \rightarrow 1$. Theoretically, contact with the wall would bring the particle to rest ( $U_{o}=0, \omega=0$ ), since an infinite force would then be required to produce motion of the sphere (Goldman et al. 1967a,b; O'Neill \& Stewartson 1967). As was discussed by Goldman et al. (1967a) such behaviour is unlikely to be observed in practice owing to a breakdown in the assumptions regarding perfect smoothness of the rigid surfaces and the integrity and constancy of physical properties of the fluid phase.

According to the method of reflexions, the pressure difference accompanying this particle sedimentation is (Brenner 1966)

$$
\begin{equation*}
\Delta P^{+} A / F^{\prime}=2\left(1-\beta^{2}\right)-\frac{4}{3} \lambda^{2}+O\left(\lambda^{3}\right) \quad(\beta \ll 1-\lambda) . \tag{4.9}
\end{equation*}
$$

The complementary perturbation expansion was evaluated by substituting into (4.6) the appropriate resistance coefficients tabulated by Goldman et al. (1967b). Both results are plotted in figure 3. As with the settling velocity, the pressure drop is a function of both $\lambda$ and $a / h$. Comparison with figure 2 suggests that the variation of pressure drop with lateral position is only partly attributable to the change of settling velocity with position. For a given sphere and tube the maximum pressure drop occurs for settling along the tube axis, $\beta=0$.

## Neutrally buoyant sphere in a Poiseuille flow

In the absence of external forces and torques the sphere is freely suspended by the fluid ( $F=0, T_{o}=0$ ). Solution of (2.13) for this case yields the following values for the translational and angular velocities of the neutrally buoyant sphere:

$$
\begin{equation*}
\frac{U_{o}}{V_{m}}=\frac{K^{s} L^{r}-K^{r} L^{s}}{\bar{K}^{t} L^{r}-K^{r} L^{t}}, \quad \frac{\omega}{\overline{V_{m}}}=\frac{K^{s} L^{t}-K^{t} L^{s}}{K^{t} L^{r}-K^{r} L^{t}} . \tag{4.10}
\end{equation*}
$$

From (2.13), the additional pressure drop resulting from the presence of such a sphere is then

$$
\begin{equation*}
\Delta P^{+} A / \mu V_{m}=P^{s}-\left(U_{o} / V_{m}\right) P^{t}-\left(\omega / V_{m}\right) P^{r} \tag{4.12}
\end{equation*}
$$

in which the appropriate expressions for the velocity ratios are those given by (4.10) and (4.11).

It is interesting to observe that the right-hand side of (4.10) is identical to that of (4.4). Indeed, the equivalence of these two equations can be demonstrated directly via use of the reciprocal theorem of $\S 2$ :

$$
\begin{equation*}
\left(\Delta P^{+} A / F\right)_{V_{m}=0, T_{o}=0}=\left(U_{o} / V_{m}\right)_{F=0, T_{o}=0} \tag{4.13}
\end{equation*}
$$

Hence, the asymptotic forms adopted by the neutrally buoyant translational velocity of (4.10) can be obtained from (4.6), (4.9) and (4.13). The remarks made above concerning figure 3 are now seen to apply to this velocity ratio $U_{o} / V_{m}$ too. Conversely, the dependence of the pressure drop upon the size and lateral position of a sedimenting sphere in a quiescent fluid can be determined experimentally from measurement of the transitional velocity of an identical, but neutrally buoyant, sphere at the same lateral position. Such information, albeit for relatively large spheres, is available from the experimental measurements of Goldsmith \& Mason (1962) of the particle velocities in a dilute suspension of neutrally buoyant spheres suspended in a Poiseuille flow. Their data for the translational velocities are replotted in figure 3 as the black circles. It would not be normally expected that the truncated asymptotic expansions for $\lambda \ll 1$ would produce accurate predictions in the upper portion of the range, $0.13<\lambda<0.53$, covered by these experiments. Nevertheless, reasonable agreement exists between the data and the method of reflexions expansion.

In figure 4 representative plots are presented for the dominant contributions to the additional pressure drop engendered by a small neutrally buoyant sphere. The dashed lines-valid when the sphere is far from the wall-were calculated from the expansion (Brenner 1966, 1970)

$$
\begin{equation*}
\Delta P^{+} R_{o} / \mu V_{m}=\frac{160}{3} \beta^{2} \lambda^{3}+O\left(\lambda^{5}\right) \quad(\beta \ll 1-\lambda) . \tag{4.14}
\end{equation*}
$$

For a given $\lambda$ the minimum pressure drop arises when the sphere moves along the tube axis, $\beta=0$. Although the upper bound of the eccentricity parameter is $\beta=1$, the maximum value of the contribution of order $\lambda^{3}$ is not $53 \cdot 3 \lambda^{3}$, as would be obtained by letting $\beta \rightarrow 1$ in (4.14). Rather, as can be seen from figure 4, the perturbation expansion for the case of a sphere close to the wall predicts the maximum value to be $226 \lambda^{3}$ - a value over four times as large as that predicted by the method of reflexions. This value is arrived at by expressing (4.12) as the expansion

$$
\begin{equation*}
\Delta P^{+} R_{o} / \mu V_{m}=4(h / a)\left[P_{2}^{s}-\left(U_{o} / G h\right) P_{2}^{t}-(\omega a / G h) P_{2}^{r}\right] \lambda^{3}+O\left(\lambda^{4}\right) \tag{4.15}
\end{equation*}
$$

in which $G=4 V_{m} / R_{o}$ is the Poiseuille shear rate at the tube wall. Numerical values for the coefficient of $\lambda^{3}$ in (4.15) were computed from the pressure drop coefficients of table 1 and the neutrally buoyant velocity ratio values of Goldman et al. (1967b). The maximum, occurring at $a / h=1$, corresponds to the


Figure 4. Additional pressure drop created by a neutrally buoyant sphere in Poiseuille flow ( $F=0, T_{0}=0$ ): ———, perturbation expansion (4.15); ----, method of reflexions expansion (4.14); $O$, sphere centre on tube axis.
additional pressure drop for a stationary sphere $\left(U_{o}=\omega=0\right)$ in contact with the tube wall.

It has been suggested by Greenstein \& Happel (1970) that the method of reflexions solution for a single sphere in a tube can be used to predict the pressure drop-flow rate relationship for the tube flow of dilute suspensions of non-hydrodynamically interacting spheres. These authors obtained an explicit asymptotic result, assuming the distribution of spheres to be uniform across the tube. The contributions (4.14) of each of the individual spheres were 'summed', by application of the equation

$$
\begin{equation*}
\left(\Delta P^{+} A\right)_{\text {total }}=2 \pi \int_{\beta=0}^{1} \Delta P^{+}(\beta) \beta d \beta \tag{4.16}
\end{equation*}
$$

Though this procedure may produce an adequate approximation for very small $\lambda$, the present work indicates that it is fundamentally improper to use the method of reflexions expansion over the entire tube cross-section. Equation (4.16) ignores the contribution of the complementary expansion (4.15) for spheres located immediately adjacent to the tube wall.

The symmetry of the resistance matrix in (2.13) is closely related to the corresponding symmetry of a similar (partitioned) matrix recently introduced by Himch (1972) for an isolated particle in an unbounded fluid subjected to a linear shear flow. Inclusion of a solid plane boundary on whose surface $\mathbf{v}=\boldsymbol{\theta}$ does not
affect the validity of Hinch's (1972) arguments. In place of the pressure drop force $\Delta P^{+} A$ in our analysis there appears a 'stresslet' $\mathbf{S}$, defined as

$$
\begin{equation*}
\mathbf{S}=\frac{1}{2} \int_{S_{p}}(\mathbf{r} d \mathbf{S} . \boldsymbol{\pi}+d \mathbf{S} . \pi \mathbf{r}) . \tag{4.17}
\end{equation*}
$$

In view of (3.22) and (3.13)-(3.14), a close correspondence clearly exists between the stresslet and the additional pressure drop force for a force-free particle,

$$
\int_{S_{p}} d \mathbf{S} \cdot \boldsymbol{\pi}=0 .
$$

This equivalence is not surprising in view of the fact that the stresslet and the term $\Delta P^{+} A V_{m}$ are both related to the additional energy dissipation in a linear shear flow.
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[^1]:    $\dagger$ See footnote on previous page.

